Systematic Strategies for Stochastic Mode Reduction

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• Averaging and Homogenisation

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- Quasi-Geostrophic Equations
- Stochastic Mode Reduction

- wide variety of problems have the common feature of multiscale possession.
- Time Scales
 - Hourly- Small scale convection
 - Monthly- Intra seasonal variability of Tropics
 - Annual- El Nino-Southern oscillation
- Previous Work
 - Reduced Linear Stochastic Models
 - Approximations made on resolved mode
 - All couplings with unresolved modes dropped and replaced by ad hoc stochastic terms of linear langevin type.

DOF have been split into resolved and unresolved modes

$$\frac{dx}{dt} = f(x, y)$$

$$\frac{dy}{dt} = g(x, y) + h(y)$$

where h(y) a non linear function of y

 Modify equation of motion of unresolved modes by representing Non-Linear self interaction terms between unresolved modes by stochastic terms.

$$h(y) \approx -\frac{\Gamma}{\epsilon}y + \frac{\sigma}{\sqrt{\epsilon}}\dot{W(t)}$$

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• Justified in coarse grained modelling on longer time scales as in climate.

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ightarrow rac{t}{\epsilon} \qquad \qquad rac{dx}{dt} = rac{f(x,y)}{\epsilon} \ rac{dy}{dt} = rac{g(x,y)}{\epsilon} - rac{\Gamma}{\epsilon^2}y + rac{\sigma}{\epsilon}W(t)$$

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• Equation of motion of unresolved mode(s) eliminated.

Averaging and Homogenisation

- averaging and homogenization-simplification of wide range of problems possessing multiple scales
- System of Linear Equations

$$\frac{du^{\epsilon}}{dt} = \mathcal{L}^{\epsilon} u^{\epsilon}$$

• averaging :

$$\mathcal{L}^{\epsilon} = rac{1}{\epsilon} \mathcal{L}_0 + \mathcal{L}_1 \qquad \qquad \mathcal{L}_i \in \mathbb{R}^{d imes a}$$

 $\mathcal{N}(\mathcal{L}_0)$ 1-Dimensional spanned by ϕ , $\mathcal{N}(\mathcal{L}_0^T)$ spanned by ψ

• seeking solutions of the form :

$$u^{\epsilon} = u_0 + \epsilon u_1 + \mathcal{O}(\epsilon^2)$$

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$$egin{aligned} \mathcal{L}_0 u_0 &= 0 & \mathcal{O}(rac{1}{\epsilon}) \ u_0 &= lpha \phi & lpha &= lpha(t) \in \mathbb{R} \ \mathcal{L}_0 u_1 &= rac{du_0}{dt} - \mathcal{L}_1 u_0 & \mathcal{O}(1) \end{aligned}$$

• Fredholm Alternative : solution for u_1 exists if and only if :

$$egin{aligned} &\langle\psi,rac{du_0}{dt}-\mathcal{L}_1u_0
angle=0\ &\ &rac{dlpha}{dt}=rac{\langle\psi,\mathcal{L}_1\phi
angle}{\langle\psi,\phi
angle}lpha \end{aligned}$$

which has non trivial solution provided

$$egin{aligned} &\langle\psi,\mathcal{L}_1\phi
angle
eq 0 \ &\lim_{\epsilon\ll1}u^\epsilon(t)pprox u_0(t)=lpha(t) \end{aligned}$$

homogenisation if

$$\langle \psi, \mathcal{L}_1 \phi \rangle = 0$$

for non trivial dynamics,

$$egin{aligned} t o rac{t}{\epsilon} \ \mathcal{L}^\epsilon &= rac{1}{\epsilon^2}\mathcal{L}_0 + rac{1}{\epsilon}\mathcal{L}_1 \end{aligned}$$

without loss of generality, we can study,

$$\mathcal{L}^{\epsilon} = rac{1}{\epsilon^2} \mathcal{L}_0 + rac{1}{\epsilon} \mathcal{L}_1 + \mathcal{L}_2$$
 $\mathcal{L}_i \in \mathbb{R}^{d imes d}$

$$u_{1} = -\alpha \eta \qquad \mathcal{L}_{0}\eta = \mathcal{L}_{1}\phi$$
$$\frac{d\alpha}{dt} = \frac{\langle \psi, \mathcal{L}_{2}\phi - \mathcal{L}_{1}\eta \rangle}{\langle \psi, \phi \rangle} \alpha$$

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Quasi-Geostrophic Equations-An Overview

• Shallow Water Equations

$$\frac{\partial \vec{u}}{\partial t} + (\vec{v} \cdot \nabla)\vec{u} + \vec{f} \times \vec{u} = -g\nabla_z h$$
$$\frac{\partial h}{\partial t} + H\nabla \cdot \vec{u} = 0$$

• Rossby No.

$$R_0 = \frac{U}{fL}$$

• $R_0 = 0 \Rightarrow$ Geostrophic Balance :

$$\vec{f} \times \vec{u} = -g\nabla_z h$$

• Quasi-Geostrophy R₀ small

$$\vec{f} imes \vec{u} \approx -g
abla_z h$$

$$\frac{D}{Dt} \left(\nabla^2 \psi + \beta y - \frac{f_0^2}{gH} \psi \right) = 0$$
$$u = U + u'$$
$$v = v'$$
$$u = \frac{\partial \psi}{\partial y}$$
$$v = -\frac{\partial \psi}{\partial x}$$
$$\psi = U\psi + \psi'$$

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- Ideal Barotropic QG equations with large scale zonal mean flow U
- $2\pi imes 2\pi$ periodic domain

$$\frac{\partial q}{\partial t} + \nabla^{\perp} \cdot \nabla q + U \frac{\partial q}{\partial x} + \beta \frac{\partial \psi}{\partial x} = 0$$
$$q = \nabla^2 \psi + h(x, y)$$
$$\frac{dU}{dt} = \frac{1}{4\pi^2} \int h \frac{\partial \psi}{\partial x} dx dy$$

• Equations expanded in Fourier modes:

$$f(x,y) = \sum_{k_x,k_y} f(\vec{k}) \exp(i(k_x x + k_y y))$$

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where f(x, y) is a 2π periodic function • truncation done at $|k|^2 \leq \Lambda$ • Truncated equations conserve energy and enstrophy

$$E_{\Lambda} = \frac{1}{2}U^{2} + \frac{1}{2}\int (\nabla\psi_{\Lambda})^{2}d\vec{x}$$
$$\frac{dE_{\Lambda}}{dt} = U_{\Lambda}\frac{dU_{\Lambda}}{dt} + \frac{d}{dt}\left[\int [\nabla \cdot (\psi_{\Lambda}\nabla\psi_{\Lambda}) - \psi_{\Lambda}\nabla^{2}\psi_{\Lambda}]d\vec{x}\right]$$
$$= U_{\Lambda}\int h_{\Lambda}\frac{\partial\psi_{\Lambda}}{\partial x}d\vec{x} + \beta\int\psi_{\Lambda}\frac{\partial\psi_{\Lambda}}{\partial x}d\vec{x} + U_{\Lambda}\int\psi_{\Lambda}\frac{\partial q_{\Lambda}}{\partial x}d\vec{x} + \int\psi_{\Lambda}(\nabla^{\perp}\psi_{\Lambda}\cdot\nabla q_{\Lambda})d\vec{x}$$
$$= U_{\Lambda}\int h_{\Lambda}\frac{\partial\psi_{\Lambda}}{\partial x}d\vec{x} + U_{\Lambda}\int\psi_{\Lambda}\frac{\partial h_{\Lambda}}{\partial x}d\vec{x}$$
$$= 0$$

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$$\epsilon_{\Lambda} = \beta U + \frac{1}{2} \int q_{\Lambda}^2 d\vec{x}$$

• The fully a priori strategy assumes that the climate PDF is Gaussian with following mean and variance for fixed α, μ .

$$\overline{U} = -\frac{\beta}{\mu}$$

$$var(U) = \frac{1}{\alpha\mu}$$

$$\overline{\psi_k} = -\frac{h_k}{\mu + |\vec{k}|^2}$$

$$var(\psi_k) = \frac{1}{\alpha |\vec{k}|^2 (\mu + |\vec{k}|^2)}$$

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Nondimensional variables for perturbations about the climate mean are given by

$$U^{new} = \frac{(U - \overline{U})}{\sqrt{var(U)}}$$
$$\psi^{new}_k = \frac{(\psi_k - \overline{\psi_k})}{\sqrt{var(\psi_k)}}$$

• The truncated equations in the new variables of each Fourier mode and U are given by :

$$\frac{\partial \psi_{k}^{new}}{\partial t} = -\frac{ik_{x}}{\sqrt{\alpha\mu}} U^{new} \psi_{k}^{new} + ik_{x} H_{k}^{'} U^{new} - i\Omega_{k}^{'} \psi_{k}^{new} + \sum_{l_{x}, l_{y}} B_{\vec{k}\vec{l}} \psi_{\vec{k}-\vec{l}} \psi_{\vec{l}} + \sum_{l_{x}, l_{y}} L_{\vec{k}\vec{l}} \psi_{\vec{l}}$$

$$\frac{dU}{dt} = 2Im \sum_{k_x, k_y} k_x H'_k \psi_k$$

where,

$$egin{aligned} H_k^{'} &= h_k \sqrt{rac{\mu}{\mid ec{k}\mid^2 (\mu + \mid ec{k}\mid^2)}} \ \Omega_k^{'} &= rac{k_xeta}{\mid ec{k}\mid^2} - \overline{U}k_x \end{aligned}$$

• Invoking the approximation of stochastic consistency:

$$\sum_{\vec{kl}} \psi_{\vec{l}} \psi_{\vec{k}-\vec{l}} + \sum_{\vec{l}} L_{\vec{kl}} \psi_{\vec{l}} \approx -\gamma_k \psi_k^{new} + \sigma_k W_k$$

$$\frac{d\psi_{k}^{new}}{dt} = ik_{x}H_{k}^{'}U^{new} - \gamma_{k}(U^{new})\psi_{k} + \sigma_{k}W_{k}$$

where

$$\gamma_k(U^{new}) = \gamma_k + i\Omega'_k + irac{k_{\chi}}{\sqrt{lpha\mu}}U^{new}$$

• ψ_k^{new} can be eliminated provided $\gamma_k(U)$ is large enough compared to other terms.

$$rac{dU}{dt} = -\gamma(U)U + \sqrt{2\gamma(U)}\xi(t)$$

where

$$\gamma(U) = 2\sum_{\vec{k}} \frac{k_x^2 | H_k' |^2 \gamma_k}{\gamma_k^2 + (\Omega_k' + k_x U(\alpha \mu)^{0.5})^2}$$

$$\xi(t) = \sqrt{\frac{2}{\gamma(U)}} Im\left(\sum_{k} k_{x} H_{k}^{'} \sigma_{k} \int_{-\infty}^{t} \exp(-\gamma_{k}(U)(t-t^{'})) \dot{W}_{k} dt^{'}\right)$$

using

$$\langle \dot{W}_{k}(s)\dot{W}_{k'}(s')\rangle = \delta(s-s')\delta_{kk'},$$

 $\xi(t)$ is delta correlated in time which implies

$$\langle \xi(t)\xi(t^{'})
angle pprox \delta(t-t^{'})$$

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Under the assumption

$$\frac{1}{(\alpha\mu)} \mid k_{x}U \mid^{2} \ll \gamma_{k}^{2} + (\Omega_{k}^{'})^{2}$$

a standard predicted linear stochastic model for U emerges :

$$rac{dU}{dt} = -2\gamma(U)U + \sqrt{2\gamma(U)}\dot{W}_k$$

where

$$\gamma(U) = \sum_{k} \frac{k_x^2 \mid H_k \mid^2 \gamma_k}{\gamma_k^2 + (\Omega'_k)^2}$$

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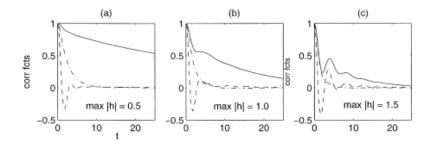


Figure: comparison of correlation function of the mean U (solid line), $Re\psi_{1,0}$ (dashed line), and $Re\psi_{0,1}$ (dot-dashed line) for different values of H

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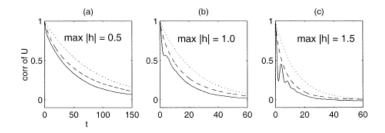


Figure: comparison of the correlation function of U DNS (solid line); nonlinear reduced stochastic model(dashed line); corresponding linear reduced stochastic model(dotted line)

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