# Seeing the invisible with black hole superradiance sourced gravitational waves

Indra Kumar Banerjee **IISER Berhampur** 

#### 06/01/2025

Hearing Beyond the Standard Model with Cosmic Sources of Gravitational Waves International Center for Theoretical Sciences





शिक्षा मंत्रालय **MINISTRY OF EDUCATION** 

## Superradiance



#### Rotating cylinder with dissipation or absorption.

## Superradiance



 $|\Omega_c>v_{\phi}|$ Absorption causes the enhancement of velocity

## Wave Superradiance



 $\Omega_c > \omega/m$ 

## **Black hole Superradiance**



 $\Omega_{
m BH} > \omega/m$ 



$$V(r)=-rac{GM_{
m BH}\mu}{r}=-rac{lpha}{r}$$

Cartoon taken from talk titled "Black holes superradiance of self-interacting scalar field" by Oliver Simon

Gravitational fine-structure Constant

This makes the massive bound states very similar to the Hydrogen atom problem with the states characterized with (n, l, m)

## **Gravitational Atom**

- Cloud size :  $r_{
  m cloud} pprox rac{n^2}{lpha^2} r_g$
- One particle energy:  $E \approx \mu \left(1 \frac{\alpha^2}{2n^2}\right) + i\Gamma_{nlm}^{\text{SR}}$

Superradiance rate 1. Most efficient at l = m2.  $\Gamma_{211}^{SR} \gg \Gamma_{322}^{SR} \gg \dots$ 

## **Gravitational Atom: Evolution**



$$ullet lpha pprox 0.05 \left( rac{M_{
m BH}}{M_{\odot}} 
ight) \left( rac{\mu}{7 imes 10^{-19} \ {
m GeV}} 
ight)$$

• 
$$t_{AB} \ll t_{BC}$$

$$ullet lpha=0.2 \hspace{0.1in} ext{and} \hspace{0.1in} \mu=6 imes10^{-22} ext{ GeV}$$

$$t_{AB} \sim 1 ~{
m yr}~~{
m and}~~t_{BC} \sim 10^6 ~{
m yr}$$

#### Gravitational Wave Generation



**Annihilation of the scalars** 

**Transition of energy levels** 

#### Gravitational Wave Generation



 $\bullet \quad {\rm GW} \ {\rm frequency} \sim 2 \mu$ 

Transition of energy levelsJCAP 11 (2024) 045• (3,1,1) 
ightarrow (2,1,1)

$$ullet \Delta \omega = rac{5}{72} lpha^2 \mu$$

$$egin{aligned} & au_{
m GW} = 0.48 \Big( rac{\mu}{10^{-21}\,{
m GeV}} \Big)^{-9} \Big( rac{M_{
m BH}}{20 M_{\odot}} \Big)^{-8} \, {
m yr} \ & E_{
m GW} = rac{lpha M_{
m BH}}{3 \mu} ig( 1 - e^{-t/ au_{
m GW}} ig) \Delta \omega \end{aligned}$$

## **Main Scheme**



# Main Scheme





JCAP 11 (2024) 045, IKB, Ujjal Kumar Dey

#### **First Order Phase Transition**

Talks by Ryusuke Jinno, Francesco Resigno, Andrew Long, Thomas Konstandin

## Important Variables: $lpha,\ eta/H,\ T_{ m reh},\ v_w,\dots$

 $\begin{aligned} & \text{GW Generation: Main contribution from bubble wall collision} & \text{Kosowsky, Turner '93} \\ & \Omega_{\text{GW}}(f) = \frac{1}{\rho_c} \frac{d\rho_{\text{GW}}}{d\ln f} \\ & = 1.67 \times 10^{-5} \left(\frac{H}{\beta}\right)^2 \left(\frac{\kappa\alpha}{1+\alpha}\right)^2 \frac{0.11 v_w^2}{0.42 + v_w^3} \left(\frac{100}{g_*}\right)^{1/3} \frac{3.8(f/f_p)^{2.8}}{1+2.8(f/f_p)^{3.8}} \\ & f_p = \frac{0.62}{1.8 - 0.1 v_w + v_w^2} \left(\frac{\beta}{H}\right) \frac{T}{100 \text{ GeV}} \left(\frac{g_*}{100}\right)^{1/6} \times 1.65 \times 10^{-5} \text{ Hz} \end{aligned}$ 

## **First Order Phase Transition**

**Origin of Primordial Black Holes**: Collapse of overdense regions<sup>1</sup>, collapse from the curvature perturbations created due to the fluctuation of bubble nucleation<sup>2</sup>, etc.

$$egin{split} \mathcal{P}_{coll} &= \exp\left[-0.5646(eta/H)^{1.466}(1+\delta_c^{0.6639eta/H})
ight] \ f_{ ext{PBH}} &= \left(rac{\mathcal{P}_{coll}}{2.2 imes10^{-8}}
ight) \left(rac{T_{ ext{reh}}}{140 ext{ MeV}}
ight) \ M_{ ext{PBH}} &= \left(rac{20}{g_*(T_{ ext{reh}})}
ight)^{1/2} \left(rac{T_{ ext{reh}}}{140 ext{ MeV}}
ight)^{-2} \end{split}$$

<sup>1</sup>2106.05637, 2212.14037, 2305.04924, 2305,04942, ... <sup>2</sup>2402.04158

## **SGWB due to BH Superradiance**

• The initial spin of PBHs created from FOPT in radiation dominated universe is very small  $\rightarrow$  we consider merger remnants with  $\chi \sim 0.7$ JHEP (2024) 006, IKB, Ujjal Kumar Dey  $f = f = f = \frac{1}{2}$  Astro. J. Lett. 704 (2009) L40

$$\begin{split} \Omega_{\rm GW}^{\rm sup} &= \frac{J}{\rho_c} \int dz \frac{dt}{dz} \int \tau p(\chi) d\chi \frac{dE_s}{df_s} \\ \tau &= \int \frac{dm_1}{m_1} \frac{dm_2}{m_2} \frac{d\tau}{d\ln m_1 d\ln m_2}, \\ \frac{d\tau}{nm_1 d\ln m_2} &= 14.8 {\rm yr}^{-1} {\rm Gpc}^{-3} f(m_1) f(m_2) \frac{(m_1 + m_2)^{10/7}}{(m_1 m_2)^{5/7}} \left(\frac{\delta_{\rm loc}}{10^8}\right) \left(\frac{v_0}{10 {\rm km/s}}\right)^{-11/7} \\ \frac{dt}{dz} &= \frac{1}{H_0 \sqrt{\Omega_{\rm M}(1+z)^3 + \Omega_{\Lambda}(1+z)}}, \\ \frac{dE_s}{df_s} &= E_{\rm GW} \delta \left(f(1+z) - f_s\right). \end{split}$$

## **SGWB** due to **BH** Superradiance

The initial spin of PBHs created from FOPT in radiation dominated universe is very small  $\rightarrow$  we consider merger remnants with  $\chi \sim 0.7$ JHEP (2024) 006, IKB, Ujjal Kumar Dev Astro. J. Lett. 704 (2009) L40  $\frac{d\tau}{d\ln m_1 d\ln m_2} = 14.8 \text{yr}^{-1} \text{Gpc}^{-3} f(m_1) f(m_2) \frac{(m_1 + m_2)^{10/7}}{(m_1 m_2)^{5/7}} \left(\frac{\delta_{\text{loc}}}{10^8}\right) \left(\frac{v_0}{10 \text{km/s}}\right)^{-11/7},$  $\frac{dt}{dz} = \frac{1}{H_0 \sqrt{\Omega_{\rm M} (1+z)^3 + \Omega_\Lambda} (1+z)},$  $\frac{dE_s}{df_s} = E_{\rm GW} \delta \left( f(1+z) - f_s \right).$ JCAP 11 (2024) 045, Phys. Dark. Univ. 36 (2022) 101009

#### **SGWB:** The two components

• The peak frequencies:  $f_p^{
m SR} = 6.843 imes 10^3 \left(rac{lpha}{0.5}
ight) \left(rac{T_{
m reh}}{
m GeV}
ight)^2 \left(rac{g_*}{100}
ight)^{1/2} 
m Hz$ 



## **SGWB: The two components**

• The 'cumulative' spectra:



## **Summary and Future Prospects**

- BH superradiance can act as a very efficient source of gravitational waves.
- Confirm or rule out the mechanism of PBH creation from FOPT or the existence of ultra-light scalars of certain mass.
- One can also consider transient monochromatic GW signals from these PBHs (memory!). Talks by Subhendra Mohanty, Ashoke Sen, Arpan Hait
- Non-standard cosmological scenarios can have implications on PBH spin which may change the spectra.

# THANK YOU